

Perturbation theory based on the physical branch of the renormalization-group equation

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A physical branch of the exact (local) renormalization-group equation which was recently found is used as a zeroth approximation for a verifiable gradient expansion.

Formally, the exact variational-derivative renormalization-group equation, which was proposed in the early studies of the fluctuation theory of phase transitions,^{1,2} makes it possible in principle to find a fixed point exactly and to calculate critical exponents. In practice, however, a perturbation theory of some sort must be used. These perturbation theories either are based on sizeable parameters^{3,4} or assume some

uncontrollable criteria for cutting off the systems of coupled equations for the vertices of the free-energy functional.⁵ The primary difficulty is that a rigorous calculation procedure should use a definitely nonlocal form of this functional.^{6,7} The expected generation of nonlocal corrections is slight to the extent that the Fisher index is small ($\eta \approx 0.03$). One can thus work with a local version of the exact renormalization-group equation as a first step⁸⁻¹² and then carry out an expansion in nonlocal corrections.

The local renormalization-group equation, however, is strongly nonlinear, and a perturbation theory (various versions of the same ϵ expansion) had to be used to solve it. The use of a φ^4 local form as a zeroth approximation,¹² for example, requires an expansion of the difference $(4 - d)$ in powers of $\eta^{1/2}$. Carrying out such an expansion is equivalent to modifying an ϵ expansion.¹³ Nevertheless, the idea of a perturbation theory based on the small extent to which nonlocal corrections are generated remains attractive.

The (unique) physical branch of the solution of the local renormalization-group equation which was recently found^{14,15} makes it possible to formulate the perturbation theory of interest in the form of a gradient expansion.¹⁶ The exact variational-derivative renormalization-group equation is¹⁷

$$\begin{aligned}
 H[\varphi] = & \frac{1}{2} \int_{\mathbf{q}} \eta(\mathbf{q}) [V - G_0^{-1}(\mathbf{q}) |\vec{\varphi}(\mathbf{q})|^2] + dV \frac{\partial H}{\partial V} \\
 & - \int d^d \mathbf{r} [(d-2) \vec{\varphi}(\mathbf{r})/2 + \mathbf{r} \nabla_{\mathbf{r}} \varphi(\mathbf{r})] \delta H / \delta \vec{\varphi}(\mathbf{r}) \\
 & + \int_{\mathbf{r} \mathbf{r}'} \{ h(\mathbf{r} - \mathbf{r}') [\delta^2 H / \delta \vec{\varphi}(\mathbf{r}) \delta \vec{\varphi}(\mathbf{r}') - \delta H / \delta \vec{\varphi}(\mathbf{r}) \delta H / \delta \vec{\varphi}(\mathbf{r}')] \\
 & - \frac{1}{2} \eta(\mathbf{r} - \mathbf{r}') \vec{\varphi}(\mathbf{r}) \delta H / \delta \vec{\varphi}(\mathbf{r}') \}, \quad (1)
 \end{aligned}$$

where $H = H_{\text{total}} - H_0$, and where H_0 is assumed to be the functional

$$\begin{aligned}
 H_0 = & \int_{\mathbf{q}} G_0^{-1}(\mathbf{q}) |\vec{\varphi}(\mathbf{q})|^2; \quad \int_{\mathbf{q}} \equiv \int d^d \mathbf{q} / (2\pi)^d; \\
 H = & \sum_{k=1}^{\infty} 2^{1-2k} \int_{\{q_i, q_{i'}\}} (2\pi)^d \delta \left(\sum_{i=1}^k (q_i + q_{i'}) \right) \times g_k \{q_i, q_{i'}\} \prod_{i=1}^k \vec{\varphi}_{q_i} \vec{\varphi}_{q_{i'}}. \quad (2)
 \end{aligned}$$

Here d is the dimensionality of the space; $\vec{\varphi}$ is an n -component vector; $h(\mathbf{q}) = \exp(-\mathbf{q}^2/2\Lambda^2)$; and the anomalous-dimensionality function $\eta(\mathbf{q})$ is given by

$$\eta(q) = \eta(0) + [(D(q) - D(0)) - \eta(0)G_0^{-1}(q)] / (G_0^{-1}(q) + g_1);$$

$$D(q) = -g_1^2 h(q) + \frac{1}{2} \int_{\mathbf{p}} h(\mathbf{p}) \{ \eta g_2(\mathbf{p}, -\mathbf{p}; \mathbf{q}, -\mathbf{q}) + g_2(\mathbf{p}, -\mathbf{q}; \mathbf{q}, -\mathbf{p}) \}. \quad (3)$$

The quantity $\eta(0) = dD/d(q^2)|_{q=0}$ is the same as the Fisher index η . Formally, at the cutoff momentum $\Lambda \rightarrow \infty$ we have $h(q) \rightarrow \text{const} = h(0) = 1$ and, correspondingly, $h(\mathbf{r} - \mathbf{r}') = [\Lambda / (2\pi)^{1/2}]^d \exp[-(\mathbf{r} - \mathbf{r}')^2 \Lambda^2 / 2] \rightarrow \delta(\mathbf{r} - \mathbf{r}')$. The solution of Eq. (1) can thus be sought in the form of a local functional $H = \Phi_0 = \int_{\mathbf{r}} f(\varphi)(\mathbf{r})$, in a procedure limited to an "ordinary" partial differential equation and $\eta = 0$:

$$\dot{f} = df - \frac{d-2}{2} \varphi \nabla_{\varphi} f + \nabla_{\varphi}^2 f - (\nabla_{\varphi} f)^2. \quad (4)$$

A local equation of this kind was studied in Refs. 14 and 15. It was found that for $2 < d < 4$ the solution of this equation has a unique physical branch which corresponds to a critical behavior..

At $d \leq 3$, an ϵ expansion in the φ^4 model is only a qualitative model of this solution. Asymptotically we have $f(\varphi) \sim \varphi^2/2$ as $\varphi \rightarrow \infty$, so a power series in $(\varphi^2)^k$ converges conditionally and cannot be truncated. On the other hand, calculations of the spectrum carried out for this solution for various values $n = 1, 2, 3, \dots$ yield good critical exponents, implying that $H = \Phi_0$ is a good zeroth approximation. The thrust of the remaining discussion is the idea of discarding the cutoff of the power series in φ in all orders of perturbation theory. This approach makes it possible to avoid an expansion in ϵ at $d \leq 3$, but it forces us to use the file $f(\varphi)$, which is known only numerically, as a zeroth approximation.

Equation (1) was, in fact, derived under the normalization $\Lambda \equiv 1$, so the substitution $H = \Phi_0$ leads to the generation of nonlocal corrections because of the term

$$- \int \tilde{h}(\mathbf{r} - \mathbf{r}') \delta H / \delta \vec{\varphi}(\mathbf{r}) \delta H / \delta \vec{\varphi}(\mathbf{r}'),$$

where

$$\tilde{h}(\mathbf{r} - \mathbf{r}') = h(\mathbf{r} - \mathbf{r}') - \delta(\mathbf{r} - \mathbf{r}').$$

Retaining the lowest nonvanishing gradients $(\nabla\varphi)^2$, and restricting this letter to the case $n = 1$, we find a generation contribution of the form $\int_{\mathbf{r}} f_{\varphi\varphi}^2 (\nabla\varphi)^2 / 2$. In this case the functional H should be supplemented: $H = \Phi_0 + \Phi_1$, where $\Phi_1 = \int_{\mathbf{r}} \chi(\varphi(\mathbf{r})) (\nabla\varphi)^2$. It is not difficult to verify that the function $\eta(\mathbf{r} - \mathbf{r}') \neq 0$ should be retained in the form $\eta(\mathbf{r} - \mathbf{r}') \approx \eta = \text{const}$ in this order. As a result, we find the following pair of equations for the functions f and χ :

$$\dot{f} = df - \frac{d-2}{2} \varphi f_{\varphi} + f_{\varphi\varphi} - f_{\varphi}^2 + B\chi;$$

$$\dot{\chi} = -[\eta + 4f_{\varphi\varphi}]\chi - \left[\frac{d-2+\eta}{2} \varphi + 2f_{\varphi} \right] \chi_{\varphi} + \chi_{\varphi\varphi} + \frac{1}{2}[f_{\varphi}^2 - \eta], \quad (5)$$

where $B = \int_p p^2 h(p) = d/(2\pi)^{d/2} \ll 1$. System of equations (5), like Eq. (4), can be integrated numerically. Since B is numerically small, and since the function χ is expected to be small in comparison with f (this point is verifiable), the last term in the equation for \dot{f} can be omitted in the lowest order in χ . We cannot say this of the parameter $\eta \neq 0$ in the form $(d-2+\eta)/2$ which is to be calculated, since it makes the primary contribution to the physically important term $\propto f_{\varphi} \varphi$ as $d \rightarrow 2$ (Ref. 14). Conditions (3), however, do not determine η completely, since in this order they reduce to the boundary conditions $\chi(0) = \chi_{\varphi}(0) = 0$. Under these restrictions, and for a fixed branch of the solution, $\dot{f} = 0$, the equation for the fixed point, $\dot{\chi}(\varphi) = 0$, has a one-parameter family of solutions $\chi(\varphi; \eta)$. Again, as in the case with $f(\varphi)$, this circumstance leads to the problem of choosing the physical branch of $\chi(\varphi; \eta)$. This choice, in fact, fixes a value $\eta \neq 0$.

The difficulty, however, is that the physical branch of $f(\varphi)$ is known only numerically, and it varies with η . To get around this difficulty, we make the substitutions $\varphi = \varphi'/(2\Delta_{\varphi})^{1/2}$ and $d_{\text{eff}} = 2d/(2-\eta)$ in the equations $\dot{f} = \dot{\chi} = 0$, where $\Delta_{\varphi} = (d-2+\eta)/2$ is the scale dimensionality of the field φ . We have

$$d_{\text{eff}} f f' / (d_{\text{eff}} - 2) - \frac{1}{2} \varphi' f_{\varphi'} + f_{\varphi' \varphi'} - f_{\varphi'}^2 = 0; \quad (6a)$$

$$-[\eta/2\Delta_{\varphi} + 4f_{\varphi' \varphi'}]\chi - \left[\frac{1}{2} \varphi' + 2f_{\varphi'} \right] \chi_{\varphi'} + \chi_{\varphi' \varphi'} + \frac{1}{2}[f_{\varphi'}^2 - \eta] = 0. \quad (6b)$$

Equation (6a) is universal, and the physical branch of $f(\varphi')$ depends unambiguously on d_{eff} . Let us analyze Eq. (6b) now. Noting that at $\varphi \geq 1$ the function $f(\varphi)$ has the asymptotic behavior¹⁴ $f(\varphi) \approx \varphi^2/2 + \dots$, we find

$$\chi_{\varphi\varphi} - (4+\eta)\chi - (4+2\Delta_{\varphi})\varphi\chi_{\varphi}/2 + (1-\eta)/2 = 0,$$

or, at $\varphi \geq 1$ (and $\chi, \chi_{\varphi}, \chi_{\varphi\varphi} \rightarrow \pm \infty$)

$$\chi_{zz} + \left(\frac{1}{2} - z \right) \chi_z - \frac{(4+\eta)}{(4+2\Delta_{\varphi})} \chi = 0, \quad (7)$$

where $z = (4+2\Delta_{\varphi})\varphi^2/4$. Equation (7) is the Kummer equation.¹⁷ It determines the asymptotic behavior $\chi(\varphi) = c(\eta)\varphi^{2\gamma} \exp[(4+2\Delta_{\varphi})\varphi^2/4]$, where $\gamma = (4-d-\eta)/[2(2+d+\eta)] > 0$. Under the given conditions $\chi(0) = \chi_{\varphi}(0) = 0$, the constant c is determined by the parameter η . Note, however, that for either sign of $c \neq 0$, and at sufficiently large $|\varphi|$, the inequality $\chi \gg f$ holds. In other words, these

solutions contradict the method by which Eq. (6b) was derived (i.e., the condition $f \gg \chi$ for all φ) and are therefore "extraneous" branches of this equation. The only possibility for satisfying the inequality $f \gg \chi$ is to choose $\chi \rightarrow \text{const} = (1 - \eta) / 2(4 + \eta)$ as $|\varphi| \rightarrow \infty$. The latter restriction unambiguously determines the Fisher index $\eta = \eta(d_{\text{eff}}) = \eta(d_{\text{eff}}(d))$. Figures 1 and 2 show the physical branches of the functions $f(\varphi)$ and $\chi(\varphi)$, respectively, in the initial part of the range of φ . The inequality $\chi(\varphi) \propto 10^{-2} f(\varphi) \ll f(\varphi)$ is obvious. Table I shows the values of $f(\varphi = 0)$, η , and d found numerically for given values of d_{eff} .

At $d \geq 3.75$, the numerical methods which we are using become less efficient, but an ϵ expansion is applicable. In the opposite limit, $d \rightarrow 2$, the higher gradients become important, and the accuracy of this calculation is again degraded. As a result, the value $\eta \approx 0.28$ turns out to be too high in comparison with the exact result of the Ising model, $\eta(d = 2; n = 1) = 0.25$.

From this point on the procedure is fairly regular. In the next order of perturbation theory, higher powers of $(\nabla\varphi)$ should be retained, as should the contribution of $B\chi$ to the equation for f . The function $\eta(\mathbf{r} - \mathbf{r}')$ should also be expanded. The coefficient of the first nonvanishing correction is fixed by the condition $\Phi_2 \ll \Phi_{1,0}$ in $H = \Phi_0 + \Phi_1 + \Phi_2 + \dots$.

We conclude with a summary of the basic features of the approach proposed here and the results.

1. The choice of the local functional $\Phi_0 = \int_{\mathbf{r}} f(\varphi(\mathbf{r}))$ as the zeroth renormalization-group approximation leads to a (unique) physical branch of the solution and to

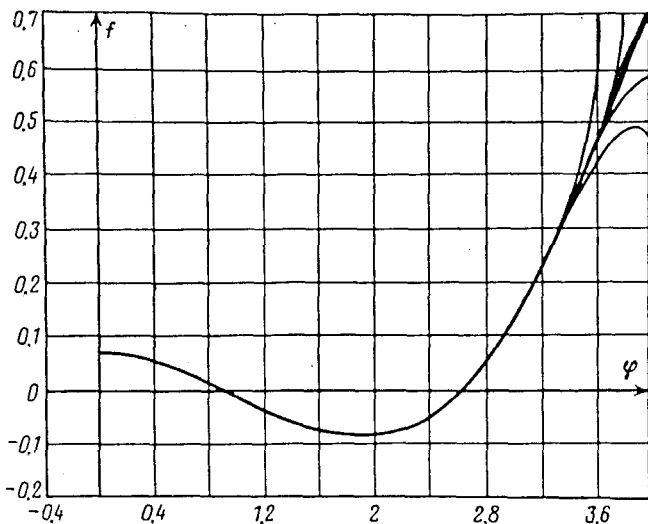


FIG. 1. Physical branch of the local renormalization-group equation (the heavy line) and nearby nonphysical branches, which diverge as $\ln|\varphi - \varphi_0|$ at finite values of φ_0 .

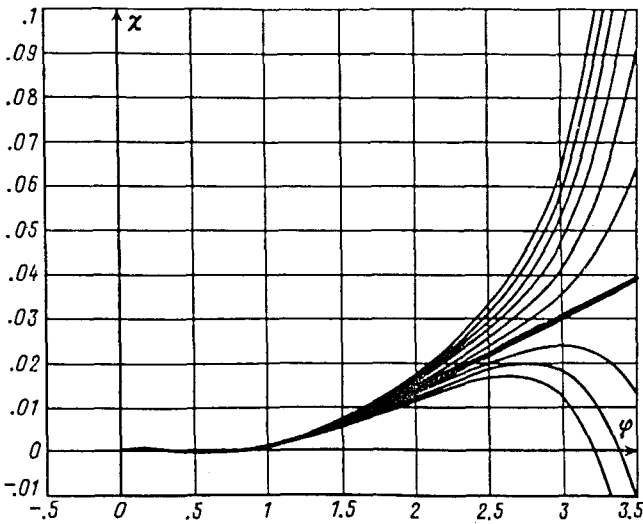


FIG. 2. The branch of $\chi(\varphi)$ which assumes a constant value (the heavy line) and side branches which diverge as $|\varphi| \rightarrow \infty$.

good critical exponents. This branch degenerates into the φ^4 model as $d \rightarrow 4$. In general, it is only qualitatively similar to the φ^4 model, it has the asymptotic behavior $\sim \varphi^2$, and it requires the retention of all powers of $(\varphi^2)^k$ in the expansion of f in powers of φ .

2. By refraining from cutting off the series in φ in all orders of perturbation theory, we are able to avoid an expansion in ϵ at $d \leq 3$, but the entire file $f(\varphi)$ found numerically must be used.

3. Incorporating deviations from a local nature leads to the appearance of corrections to the local-approximation equation, so system of equations (5) must be solved in a self-consistent fashion. Consequently, the quantity η in the combination $(d - 2 + \eta)/2$ is not comparable to the difference $(d - 2)$, which is small as $d \rightarrow 2$; on

TABLE I.

d_{eff}	$f(\varphi=0)$	η	d
2.3	0,44(3)	0,279(6)	≈ 2
2.5	0,23(1)	0,15(8)	2,30
2,75	0,13(7)	0,07(5)	2,65
3	0,076(2)	0,03(4)	2,95
3,05	0,067(7)	0,03(1)	≈ 3
3,25	0,041(1)	0,013(4)	3,23
3,5	0,020(0)	0,004(1)	3,49

the contrary, it increases the effective dimensionality d_{eff} , so the expansion works satisfactorily even for $d \rightarrow 2$.

4. The small parameter of the theory is actually the quantity $f(0)$, which can be shown to fall off exponentially with increasing d : $f(0) \sim \exp[-a(d-2+\eta)]$, where $a > 0$ is on the order of unity. On the other hand, we have $f(0) \sim \epsilon \rightarrow 0$ as $d \rightarrow 4$. In this limit the efficiency of the numerical methods which we are using is degraded, but an ϵ expansion works. As a result, the approach is most effective for $d \simeq 3$, i.e., just where we would most like it to be.

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