

Hybridization of collective oscillations in ${}^3\text{He-A}$ near the transition to the A_1 phase

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Hydrodynamic oscillations of longitudinal magnetization and entropy in a superfluid ${}^3\text{He-A}$ in a magnetic field are analyzed. It is shown that in the neighborhood of a transition to the A_1 phase there is a strong mixing of the entropy (and temperature) oscillations with the spin wave.

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Liquid ${}^3\text{He-A}$ is a mixture of two, weakly interacting superfluid components characterized by Cooper pairing in the spin states $\uparrow\uparrow$ and $\downarrow\downarrow$. In the presence of a strong magnetic field and in the immediate vicinity of the transition to a normal phase, the quasi-particle pairing amplitudes with a projection of the total spin $S_z = \pm 1$ are not equal to each other ($\Delta_+ \neq \Delta_-$). Taking into account this fact, the order parameter describing the superfield ${}^3\text{He-A}$ can be represented in the form

$$A_{\mu i} = \Delta(T) d_{\mu} u_i = \Delta(T) (\alpha_+ d_1 + i \alpha_- d_2)_{\mu} (u_1 + i u_2)_i, \quad (1)$$

where (d_1, d_2) and (u_1, u_2) are unit vectors in the spin orbital spaces, respectively, $\alpha_{\pm} = (\Delta_{\pm} / \Delta)$, and $\Delta^2 = \frac{1}{2}(\Delta_+^2 + \Delta_-^2)$.

In the hydrodynamic description of superfluid ${}^3\text{He-A}$ the energy density ϵ , in addition to the mass density ρ , entropy density $S = \rho\sigma$, momentum density \mathbf{g} , and magnetization \mathbf{M} , depends on two superfluid velocities

$$\mathbf{v}_s = (\hbar/2m)\mathbf{u}_{1i} \nabla u_{2i}, \quad \mathbf{w}_{SP} = (\hbar/2m)d_{i\mu} \nabla d_{2\mu}, \quad (2)$$

where the spin velocity \mathbf{w}_{SP} , in contrast to \mathbf{v}_s , is a Galilean invariant. Below we shall examine a linear, dissipation-free hydrodynamics for specified orientations of the axes $s = [\mathbf{d}_1, \mathbf{d}_2]$ and $l = [\mathbf{u}_1, \mathbf{u}_2]$, when

$$d\epsilon = \mu d\rho + T dS + v_n d\mathbf{g} + \omega d\mathbf{M} + \mathbf{g}_S d\mathbf{v}_S + \mathbf{g}_{SP} d\mathbf{w}_{SP}, \quad (3)$$

where \mathbf{M} is the longitudinal magnetization and $\mathbf{g} = \rho\mathbf{v}_n + \mathbf{g}_S$ because of the Galilean covariance. The question of whether the unit vector s can be included in the hydrodynamic variables was discussed elsewhere.¹

Using the well-known scheme of Ref. 2, we can write the system of questions

$$\rho + \nabla(\rho v_n + \mathbf{g}_S) = 0, \quad \mathbf{g} + \nabla P = 0, \quad (4)$$

$$S + \nabla(S v_n) = 0, \quad \mathbf{M} + \nabla(\mathbf{M} v_n + M_S \mathbf{g}_{SP}/\rho) = 0,$$

which is supplemented by the equations of motion

$$\dot{\mathbf{v}}_S = -(1/\rho) \nabla P + \sigma \nabla T + (1/\rho) \mathbf{M} \nabla \omega, \quad (5)$$

$$\mathbf{w}_{SP} = -(\gamma \hbar/2m) \nabla \omega = -(1/\rho) M_S \nabla \omega,$$

where $M_S = \gamma \hbar \rho/2m$ is the magnetization of a totally polarized liquid ${}^3\text{He}$.

As a result of splitting the states of the Cooper pairs in a magnetic field, the spin and spatial degrees of freedom of the A phase become mixed,² which also occurs in the structure of superfluid fluxes of mass \mathbf{g}_S and spin \mathbf{g}_{SP} :

$$\mathbf{g}_S = \rho_S [(\mathbf{v}_S - \mathbf{v}_n) + \alpha \mathbf{w}_{SP}],$$

$$\mathbf{g}_{SP} = \rho_S [\mathbf{w}_{SP} + \alpha(\mathbf{v}_S - \mathbf{v}_n)], \quad (6)$$

where the coefficient $\alpha = 2\alpha_+ \alpha_- = (\Delta_+^2 - \Delta_-^2)/(\Delta_+^2 + \Delta_-^2)$ characterizes the proximity of a transition to the A_1 phase for which $\Delta_+ \neq 0$ and $\Delta_- = 0$ (or vice versa). We shall not introduce in explicit form the orbital anisotropy characteristic of ${}^3\text{He-A}$, which can be easily taken into account, but it is not important for our purpose.

At some distance from a transition to the A_1 phase (where $|\alpha| \ll 1$) the superfluid ${}^3\text{He-A}$ has weakly coupled hydrodynamic modes such as first and second sound and also a (longitudinal) spin wave with propagation rates $c_1 \gg c_S \gg c_2$. As we approach the $A \rightarrow A_1$ transition, a kinematic mixing of the spin and orbital degrees of freedom arises (see Ref. 6), which leads to curious dynamic effects. As shown in Ref. 3, a strong mixing of the density and longitudinal magnetization oscillations in the fourth sound regime ($v_n \equiv 0$) occurs in the immediate vicinity of a transition to the A_1 phase. On the

other hand, if the motion of the normal component is not blocked, then we can easily obtain a system of linearized equations from Eqs. (4), (5), and (6) describing the oscillations ρ , σ , and $\xi = M/M_S$:

$$\begin{aligned} \ddot{\rho} - (\partial P / \partial \rho) \nabla^2 \rho - (\partial P / \partial \sigma) \nabla^2 \sigma - (\partial P / \partial \xi) \nabla^2 \xi &= 0, \\ \ddot{\sigma} - (\rho_S / \rho_n) \sigma^2 \nabla^2 T - (\rho_S / \rho_n) \sigma (\xi - \alpha) (M_S / \rho) \nabla^2 \omega &= 0, \end{aligned} \quad (7)$$

$$\ddot{\xi} - (\rho_S / \rho_n) [(\xi - \alpha)^2 + (1 - \alpha^2)(\rho_n / \rho)] (M_S / \rho) \nabla^2 \omega - (\rho_S / \rho_n) \sigma (\xi - \alpha) \nabla^2 T = 0.$$

If the thermal expansion and magnetostriction are disregarded, then the oscillations of the entropy and of the longitudinal magnetization are not coupled to the density oscillations and, setting $\rho = \text{const}$, we obtain closed set of equations for σ and ξ :

$$\begin{aligned} \ddot{\sigma} - u_2^2 \nabla^2 \sigma - \sigma (\xi - \alpha)^{-1} [u_5^2 - (1 - \alpha^2)(\rho_S / \rho)(M_S / \rho)(\partial \omega / \partial \xi)] \nabla^2 \xi &= 0, \\ \ddot{\xi} - u_5^2 \nabla^2 \xi - (\xi - \alpha) \sigma^{-1} [u_2^2 + (1 - \alpha^2)(\rho_S / \rho) \sigma / (\xi - \alpha) (\partial T / \partial \xi)] \nabla^2 \sigma &= 0, \end{aligned} \quad (8)$$

where

$$\begin{aligned} u_2^2(\alpha) &= (\rho_S / \rho_n) [\sigma^2 (\partial T / \partial \sigma) + \sigma (\xi - \alpha) (\partial T / \partial \xi)], \\ u_5^2(\alpha) &= (\rho_S / \rho_n) \{ [(\xi - \alpha)^2 + (1 - \alpha^2)(\rho_n / \rho)] (M_S / \rho) (\partial \omega / \partial \xi) \\ &\quad + \sigma (\xi - \alpha) (\partial T / \partial \xi) \}. \end{aligned} \quad (9)$$

Equations (8) describe two hydrodynamic modes of the split A phase of liquid ^3He , whose propagation velocity squared is given by

$$c_{\pm}^2(\alpha) = \frac{1}{2} \{ u^2(\alpha) \pm [u^4(\alpha) - 4(1 - \alpha^2)v^4]^{1/2} \}, \quad (10)$$

where

$$u^2(\alpha) = u_5^2(\alpha) + u_2^2(\alpha),$$

$$v^4 = (\rho_S / \rho) (\rho_S / \rho_n) \sigma^2 [(M_S / \rho) (\partial \omega / \partial \xi) (\partial T / \partial \sigma) - (\partial T / \partial \xi)^2].$$

At some distance from the A phase ($|\alpha| \ll 1$) there is only a weak coupling of the entropy and longitudinal magnetization oscillations (note that $M \ll M_S$). As we approach the temperature for transition to the A phase, $\alpha^2 \rightarrow 1$, and the longitudinal magnetization wave sets up an intensive oscillatory motion of entropy and temperature, dictating a high propagation velocity for them. The high-frequency branch in the A_1 phase links up with the hybrid, oscillatory mode examined recently.⁴ However, the frequency of the lower branch vanishes together with $(1 - \alpha^2)$.

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