

Neutral current with allowance for the growth of the sea of quarks and antiquarks

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It is shown that allowance for the contribution of the sea of quarks-antiquarks above the threshold of the production of charmed particles has a substantial effect in neutrino processes with neutral currents.

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Experiments at high energy point to violation of scaling in reactions of neutrino and antineutrino scattering by nucleons,^[1–4] and also in μp scattering.^[5] It was indicated in earlier papers^[6] that if this violation is interpreted as a consequence of an increase in the role of the sea quarks, then the GIM scheme explains in natural fashion the observed violation of charge symmetry^[1,2] and all the characteristics of dilepton events.^[8–10] We show in this article that the same increase of the sea has a substantial effect in neutrino processes with neutral currents.

For processes with neutral currents the neutrino interaction Lagrangian in the GIM scheme takes the form

$$L = \frac{G}{\sqrt{2}} [\bar{\nu} \gamma_{\alpha} (1 + \gamma_5) \nu] \{ \alpha [\bar{\psi}_u O_{\alpha}^{+} \psi_u + \bar{\psi}_c O_{\alpha}^{+} \psi_c] + \beta [\bar{\psi}_d O_{\alpha}^{+} \psi_d + \bar{\psi}_s O_{\alpha}^{+} \psi_s] \\ + \gamma [\bar{\psi}_u O_{\alpha}^{-} \psi_u + \bar{\psi}_c O_{\alpha}^{-} \psi_c] + \delta [\bar{\psi}_d O_{\alpha}^{-} \psi_d + \bar{\psi}_s O_{\alpha}^{-} \psi_s] \}; \quad O_{\alpha}^{\pm} = \gamma_{\alpha} (1 \pm \gamma_5).$$

In the Weinberg–Salam model we have

$$\alpha = \frac{1}{2} - \frac{2}{3} \sin^2 \theta_W, \quad \beta = -\frac{1}{2} + \frac{1}{3} \sin^2 \theta_W, \quad \gamma = -\frac{2}{3} \sin^2 \theta_W, \quad \delta = \frac{1}{3} \sin^2 \theta_W.$$

For an isoscalar target, neglecting the sea of charmed quarks and under the standard assumptions concerning the properties of the sea and valent quarks ($u = u_V + u_S$, $d = d_V + d_S$, $u_S = \bar{u}_S$, $d_S = \bar{d}_S$, $S = S_S = \bar{S}_S$) we have the following expressions for the scattering cross sections:

$$\frac{d^2\sigma(\nu N \rightarrow \nu X)}{dx dy} = \sigma_0 x \{ [a^2 + \beta^2 + (\gamma^2 + \delta^2)(1 - \gamma)^2](u + d)_V + (2 - 2\gamma + \gamma^2)[(a^2 + \beta^2 + \gamma^2 + \delta^2)(u + d)_S + 2(\beta^2 + \delta^2)s] \};$$

$$\frac{d^2\sigma(\bar{\nu} N \rightarrow \nu X)}{dx dy} = \sigma_0 x \{ [\gamma^2 + \delta^2 + (a^2 + \beta^2)(1 - \gamma)^2](u + d)_V + (2 - 2\gamma + \gamma^2)[(a^2 + \beta^2 + \gamma^2 + \delta^2)(u + d)_S + 2(\beta^2 + \delta^2)s] \}.$$

Here x and y are the usual scaling variables; u , d , s , and c are the distribution functions of the quarks in the proton, and $\sigma_0 = G^2 ME/\pi$.

Under the same assumptions, for processes with charged currents, at energy above the threshold of charmed-particle production, the cross sections take the form

$$\frac{d^2\sigma(\nu N \rightarrow \mu X)}{dx dy} = \sigma_0 x [(u + d)\cos^2\theta + 2s\sin^2\theta + (\bar{u} + \bar{d})(1 - \gamma)^2 + \underline{(u + d)\sin^2\theta + 2s\cos^2\theta}];$$

$$\frac{d^2\sigma(\bar{\nu} N \rightarrow \bar{\mu} X)}{dx dy} = \sigma_0 x [(u + d)(1 - \gamma)^2 + (\bar{u} + \bar{d})\cos^2\theta + 2\bar{s}\sin^2\theta + \underline{(\bar{u} + \bar{d})\sin^2\theta + 2\bar{s}\cos^2\theta}].$$

The underscored terms correspond to production of charmed particles, and θ is the Cabibbo angle. The total cross sections of the reactions are

$$\sigma(\nu \rightarrow \nu) = \sigma_0 \langle u + d \rangle_V [a^2 + \beta^2 + \frac{1}{3}(\gamma^2 + \delta^2) + \frac{4}{3}(a^2 + \beta^2 + \gamma^2 + \delta^2)\epsilon + \frac{4}{3}(\beta^2 + \delta^2)\lambda\epsilon];$$

$$\sigma(\bar{\nu} \rightarrow \bar{\nu}) = \sigma_0 \langle u + d \rangle_V [\gamma^2 + \delta^2 + \frac{1}{3}(a^2 + \beta^2) + \frac{4}{3}(a^2 + \beta^2 + \gamma^2 + \delta^2)\epsilon + \frac{4}{3}(\beta^2 + \delta^2)\lambda\epsilon];$$

$$\sigma(\nu \rightarrow \mu) = \sigma_0 \langle u + d \rangle_V [(1 + \epsilon)\cos^2\theta + \lambda\epsilon\sin^2\theta + \frac{\epsilon}{3} + \underline{(1 + \epsilon)\sin^2\theta + \lambda\epsilon\cos^2\theta}];$$

$$\sigma(\bar{\nu} \rightarrow \bar{\mu}) = \sigma_0 \langle u + d \rangle_V \left[\frac{1}{3}(1 + \epsilon) + \epsilon\cos^2\theta + \lambda\epsilon\sin^2\theta + \underline{\epsilon\sin^2\theta + \lambda\epsilon\cos^2\theta} \right];$$

$$\epsilon = \langle u + d \rangle_S / \langle u + d \rangle_V, \quad \lambda = 2\langle s \rangle / \langle u + d \rangle_S, \quad \langle u \rangle = \int x u dx.$$

In the usual quark-parton model, the distribution functions u , d , s and c are assumed to depend only on the variable x . Violation of scaling means that these functions can depend also on q^2 . We assume that the dependence on q^2 is weak (for example, in gauge theories with asymptotic freedom this dependence

is logarithmic). Therefore, when integrating the expression for $d^2\sigma/dx dy$ we can assume that the distribution functions are taken at a certain effective q^2 that depends on the energy.

In the absence of a sea of quarks and below the threshold of charmed-particle production we have $R_\nu = R_\nu^0 = [\alpha^2 + \beta^2 + \frac{1}{3}(\gamma^2 + \delta^2)]/\cos^2\theta$, $R_{\bar{\nu}} = R_{\bar{\nu}}^0 = 3(\gamma^2 + \delta^2) + \alpha^2 + \beta^2$. In the Weinberg—Salam model at $\sin^2\theta_W = 3/8$,

$$R_\nu^0 = 11/(48\cos^2\theta) \approx 0.24; \quad R_{\bar{\nu}}^0 = 7/16 \approx 0.44 \quad .$$

In the region above the threshold of charmed-particle production and at $\epsilon(E) \neq 0$ we have

$$R_\nu = [\alpha^2 + \beta^2 + \frac{1}{3}(\gamma^2 + \delta^2) + \frac{4}{3}(\alpha^2 + \beta^2 + \gamma^2 + \delta^2)\epsilon + \frac{4}{3}(\beta^2 + \delta^2)\lambda\epsilon]/(1 + \frac{4}{3}\epsilon + \lambda\epsilon);$$

$$R_{\bar{\nu}} = [\gamma^2 + \delta^2 + \frac{1}{3}(\alpha^2 + \beta^2) + \frac{4}{3}(\alpha^2 + \beta^2 + \gamma^2 + \delta^2)\epsilon + \frac{4}{3}(\beta^2 + \delta^2)\lambda\epsilon]/(\frac{1}{3} + \frac{4}{3}\epsilon + \lambda\epsilon).$$

At $\sin^2\theta_W = 3/8$

$$R_\nu = \frac{1}{16}(11 + 18\epsilon + 10\lambda\epsilon)/[3 + \epsilon(4 + 3\lambda)];$$

$$R_{\bar{\nu}} = \frac{1}{16}(7 + 18\epsilon + 10\lambda\epsilon)/[3 + \epsilon(4 + 3\lambda)].$$

At the parameter values $\lambda = 0.6$ and $\epsilon = 0.25$ ^[6] we get

$$R_\nu = 0.24; \quad R_{\bar{\nu}} \approx 0.32.$$

In the limit as $\epsilon \rightarrow \infty$,

$$R_\nu = R_{\bar{\nu}} = R^\infty = \frac{1}{4} + \frac{1}{8}(1 - \lambda)/(4 + 3\lambda).$$

At $\lambda = 0.6$ the quantity R^∞ is equal to $R^\infty \approx 0.26$.

Thus, the increase of the parameter ϵ with energy, obtained from processes with charged currents,^[1-4] leads to a substantial change in the value of $R_{\bar{\nu}}$ with increasing energy (the value of R_ν remains practically unchanged in this case).

The ratio of the cross sections for the neutral currents in the Weinberg—Salam model is equal to (at $\sin^2\theta_W = 3/8$)

$$\sigma_{\bar{\nu} \rightarrow \bar{\nu}}/\sigma_{\nu \rightarrow \nu} = [7 + 2\epsilon(9 + 5\lambda)]/[11 + 2\epsilon(9 + 5\lambda)].$$

With increasing energy, this ratio can range between $7/11 \approx 0.64$ and unity.

We note also the following circumstance. When processes with charged and neutral currents are separated in experiment, a kinematic cutoff in the hadron energy is usually employed. The true values R_ν and $R_{\bar{\nu}}$ are then reconstructed by using a definite model-dependence expression for $d^2\sigma/dx dy$. It is clear that at high energies such a correction must be carried out with allowance for the increase in the fraction of the sea quarks (see the formulas for $d^2\sigma/dx dy$).

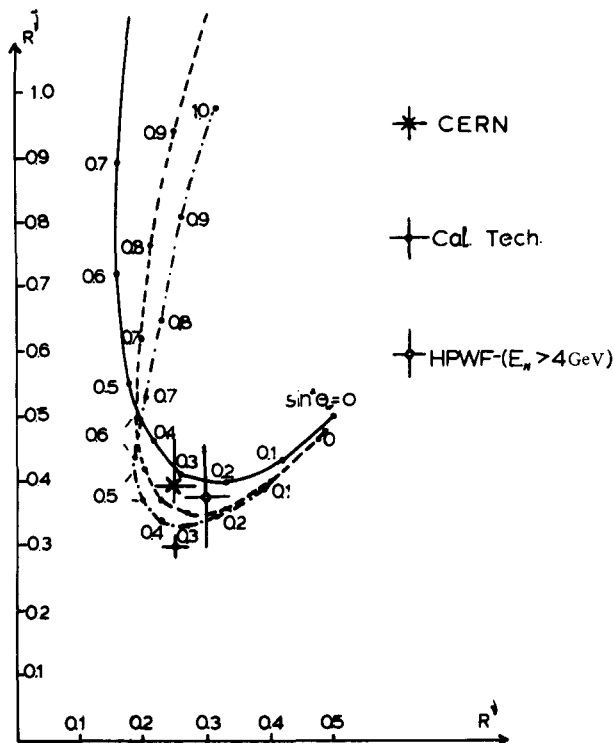


FIG. 1.

Figure 1 shows the results of a comparison of the experimental data with the Weinberg—Salam model. It is seen that allowance for the sea of quarks and antiquarks improves the agreement.

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